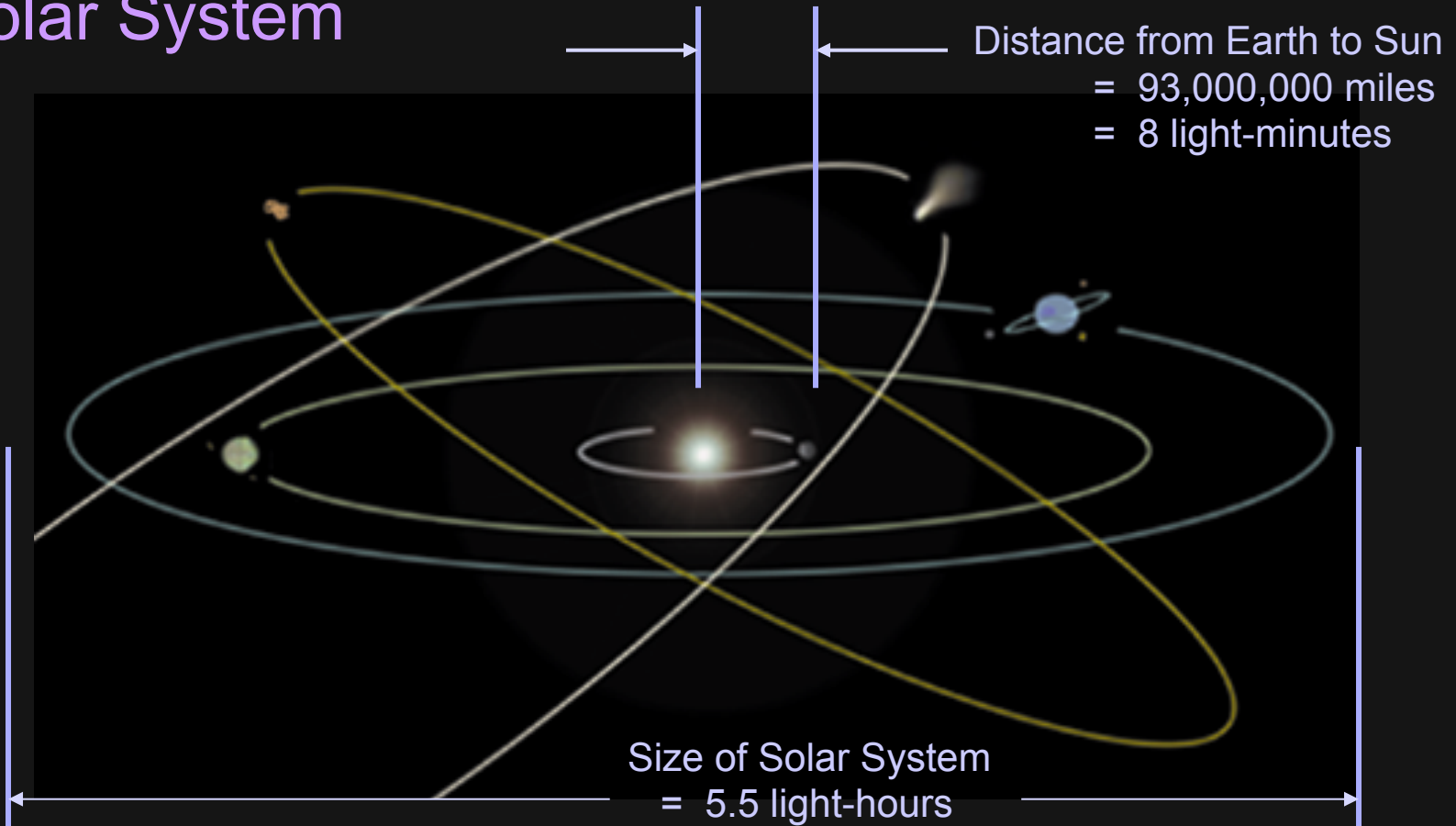


What is a Galaxy ?

Solar System



What is a Galaxy?

Stellar Region

Sun
(solar system
too small to be
seen on this scale)

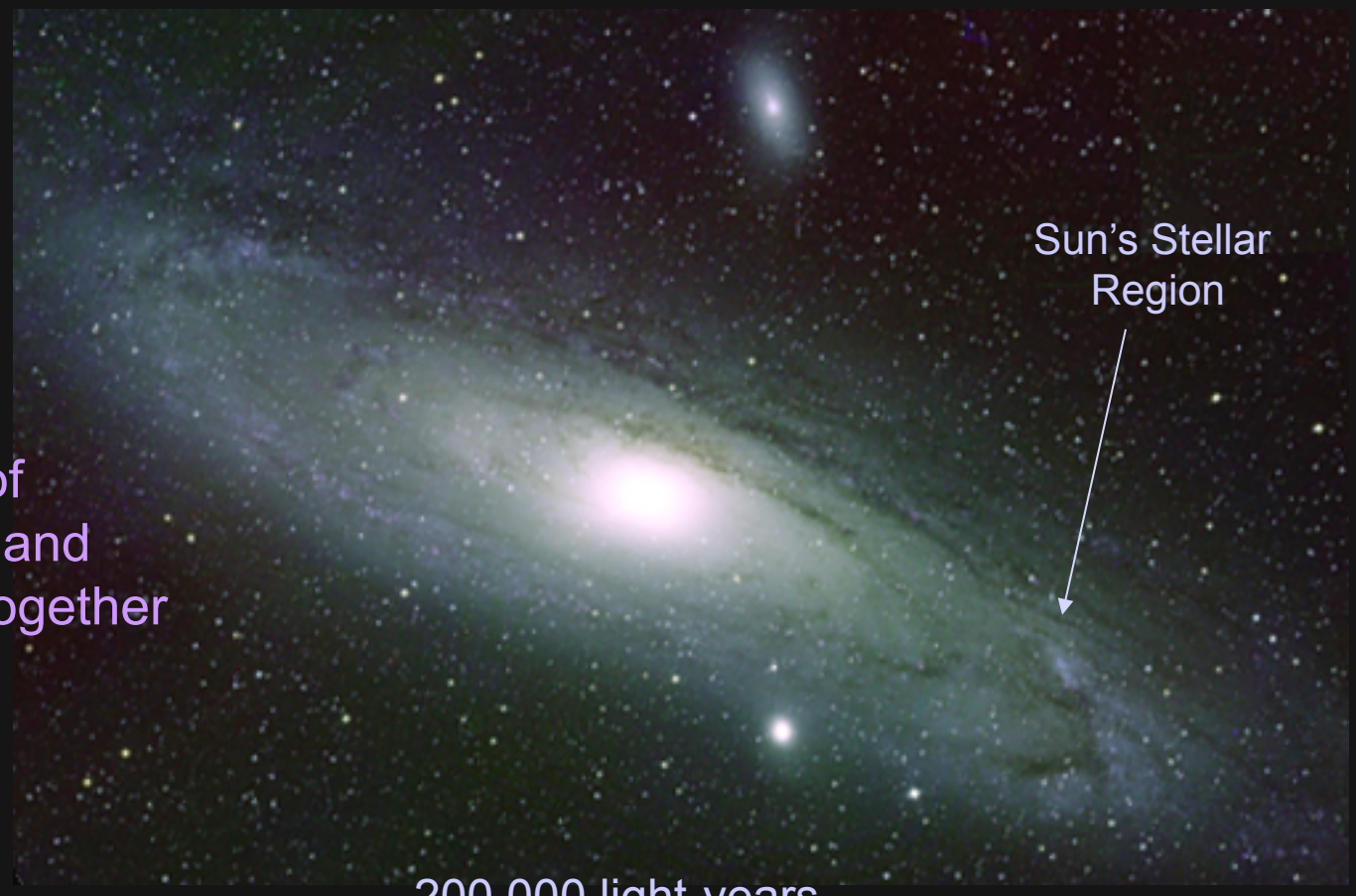


30
light-years

What is a Galaxy?

Galaxy

a massive collection of stars, gas, and dust kept together by gravity



200,000 light-years

What is a Galaxy?



**If our solar system
was the size
of a cell in
the human body,
our galaxy
would still
measure over
one mile across.**

Types of Galaxies

Spiral

disk-like

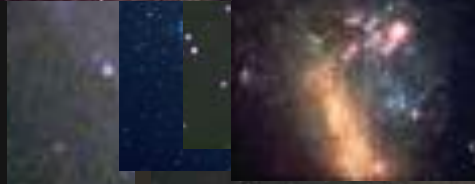
with arms of stars and dust forming a spiral pattern.

Barred spiral galaxies have a bright central bar of stars and dust through which the spiral arms pass.

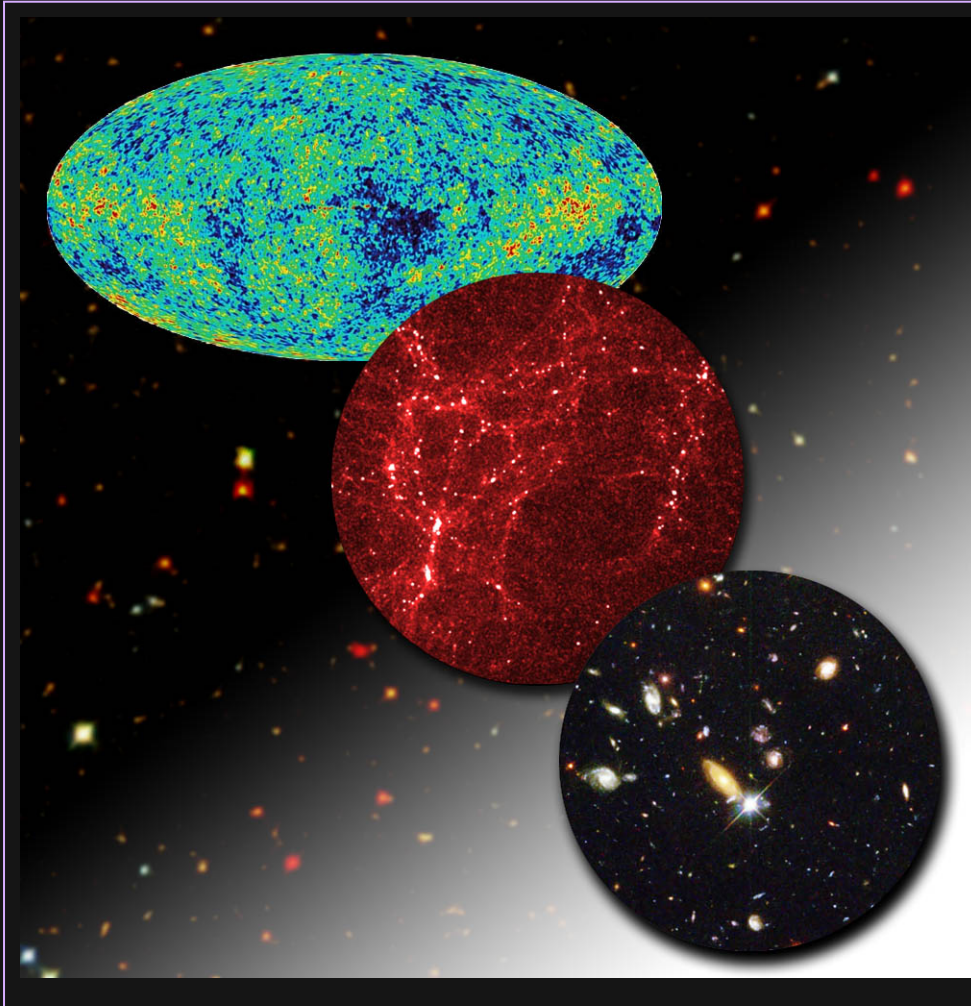
Elliptical galaxies are ellipsoidally-shaped, with less gas and dust than spiral galaxies.

They range in shape from "spheroidal" to "rod-like".

Peculiar galaxies are distorted or irregular in shape, often defined as galaxies that have been distorted by a collision with another galaxy or similar catastrophe.



Galaxy Formation



Galaxies form from the primordial density fluctuations that arise after the big bang and grow under inflation.

These density fluctuations form filaments, and galaxies form in knots along the filaments.

Spirals vs. Ellipticals

Final type of galaxy depends on initial rate of star formation:

- If stars form quickly, then galaxy becomes elliptical. Stars form within initial distribution of gas, and follow their initial orbits.
- If stars form later, the gas has time to collapse into a disk. Most stars form within the disk. The galaxy becomes a spiral.

Formation via Galaxy Mergers

In clusters, galaxies can pass close to one another.

- Galaxies can become distorted, and often merge.
- Mergers often lead to giant elliptical galaxies at the heart of large clusters.

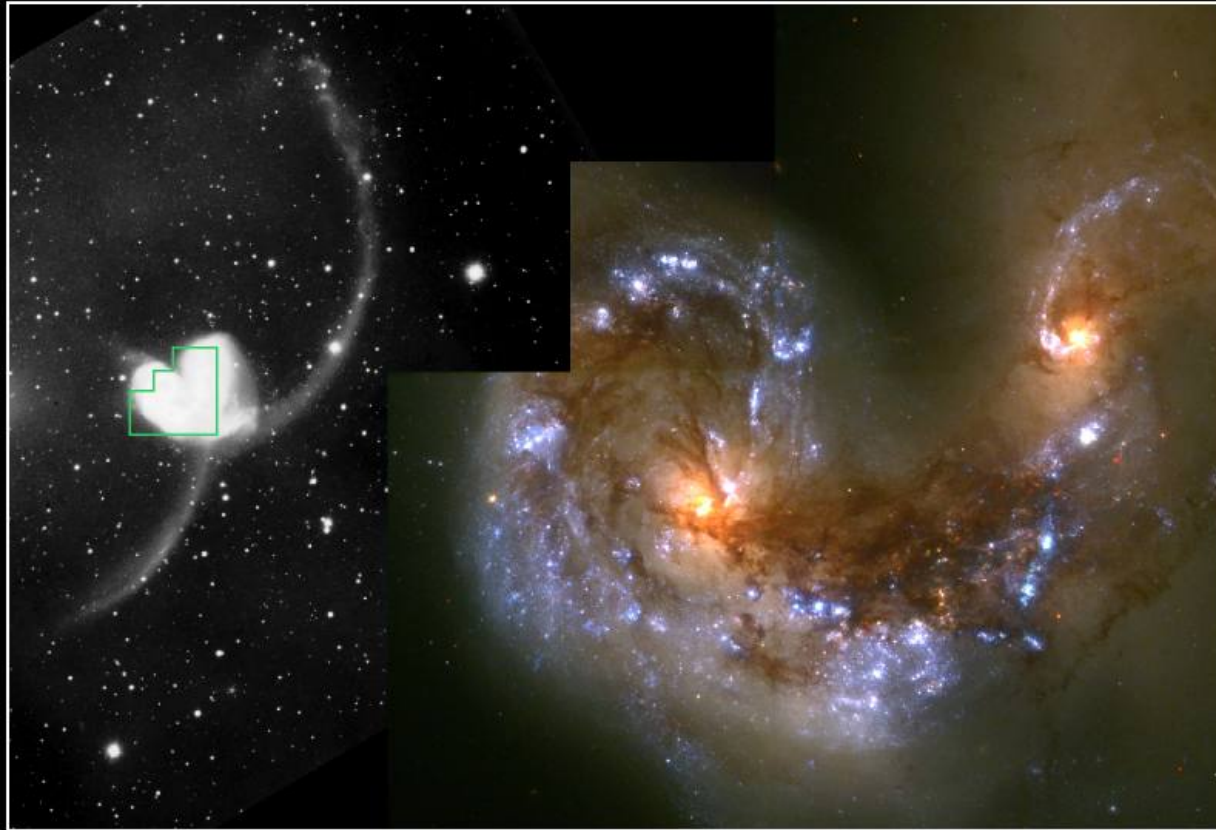
Spirals in Grazing Encounter

Galaxies NGC 2207 and IC 2163



Hubble
Heritage

Antennae Galaxies



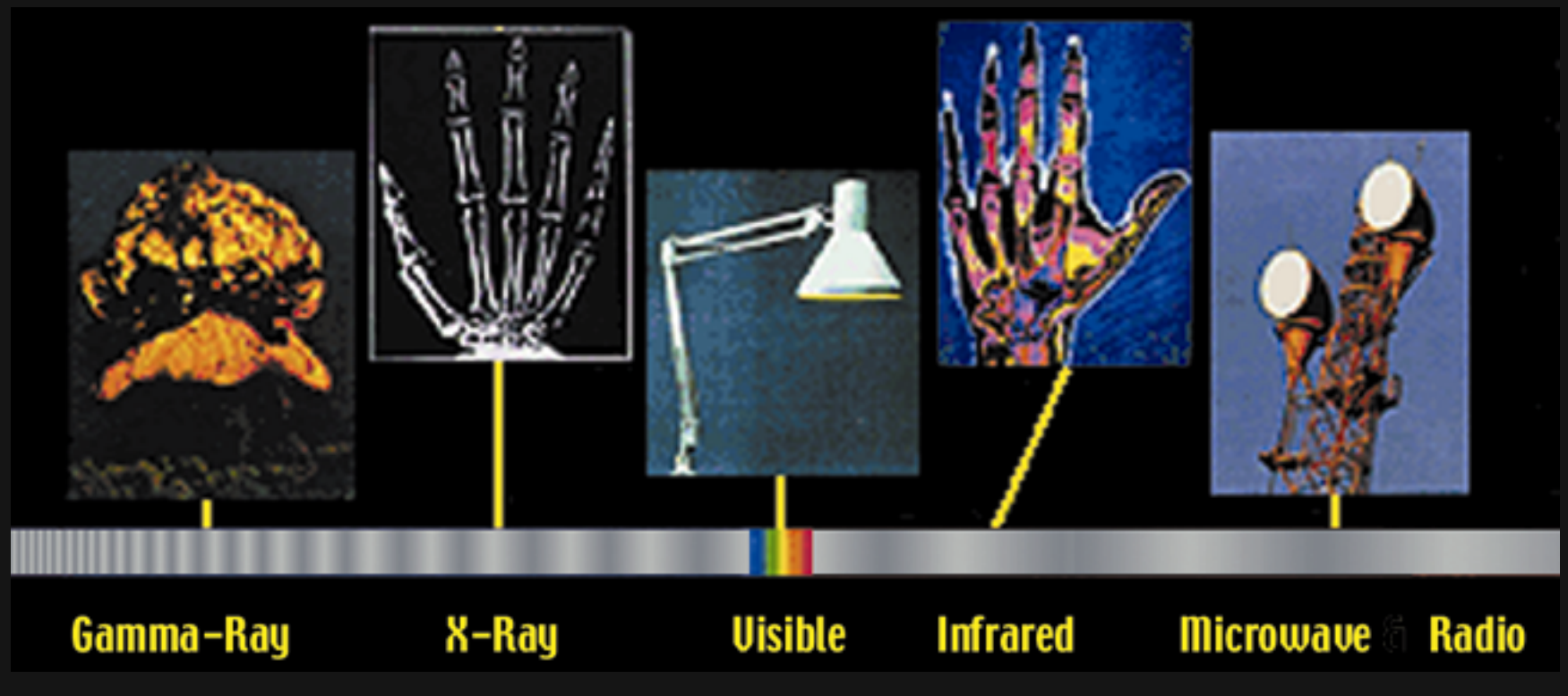
Colliding Galaxies NGC 4038 and NGC 4039

HST • WFPC2

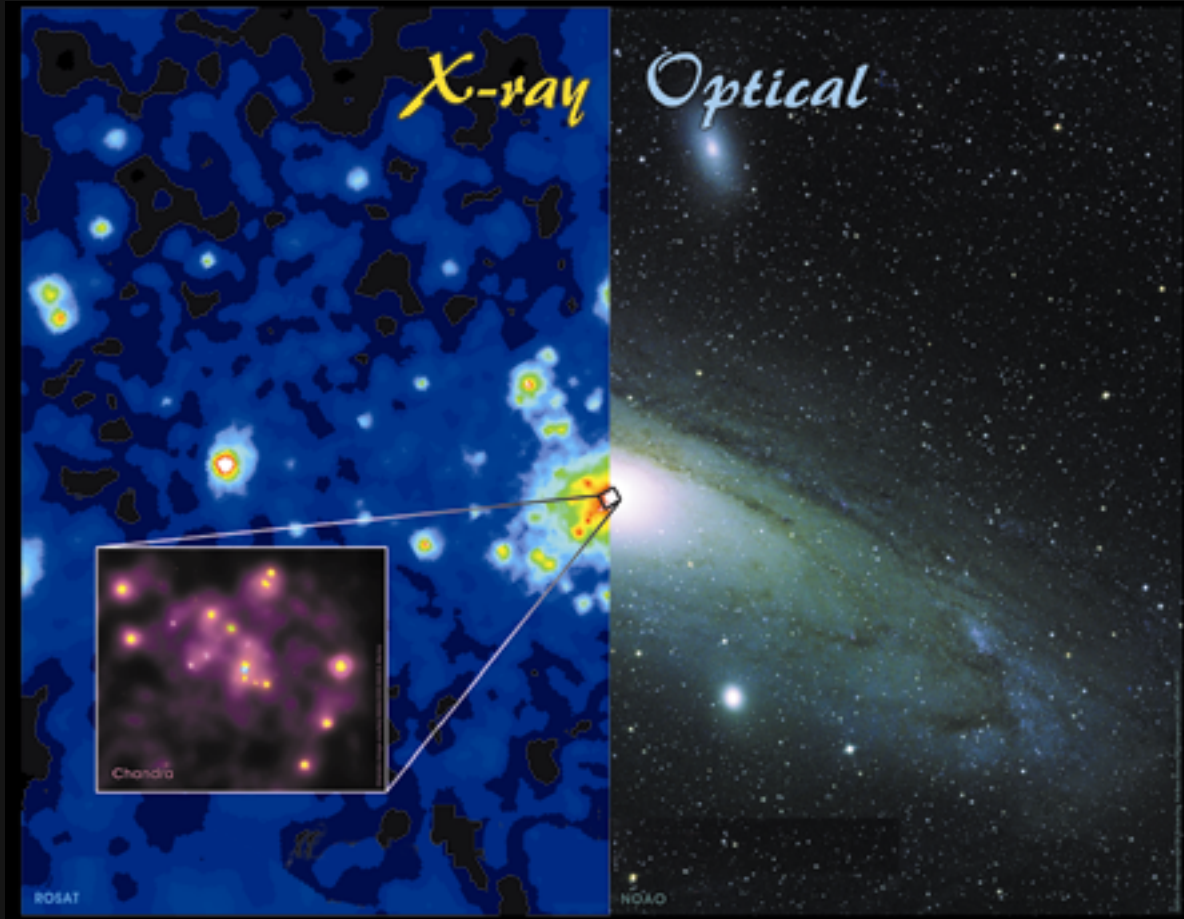
PRC97-34a • ST ScI OPO • October 21, 1997 • B, Whitmore (ST ScI) and NASA

“Invisible” Light from Galaxies

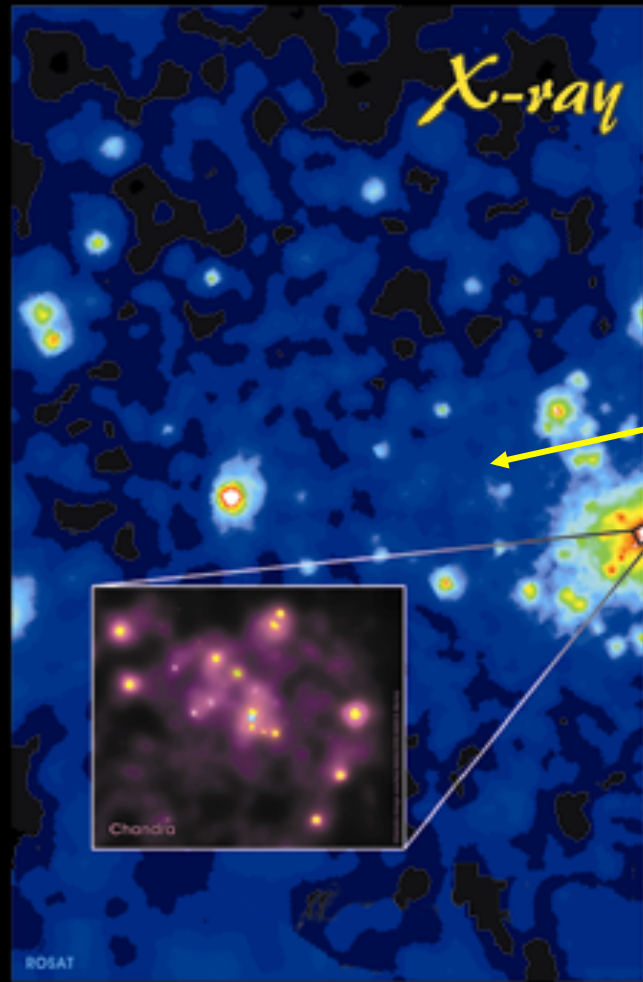
Electromagnetic Spectrum



“Invisible” Light from Galaxies

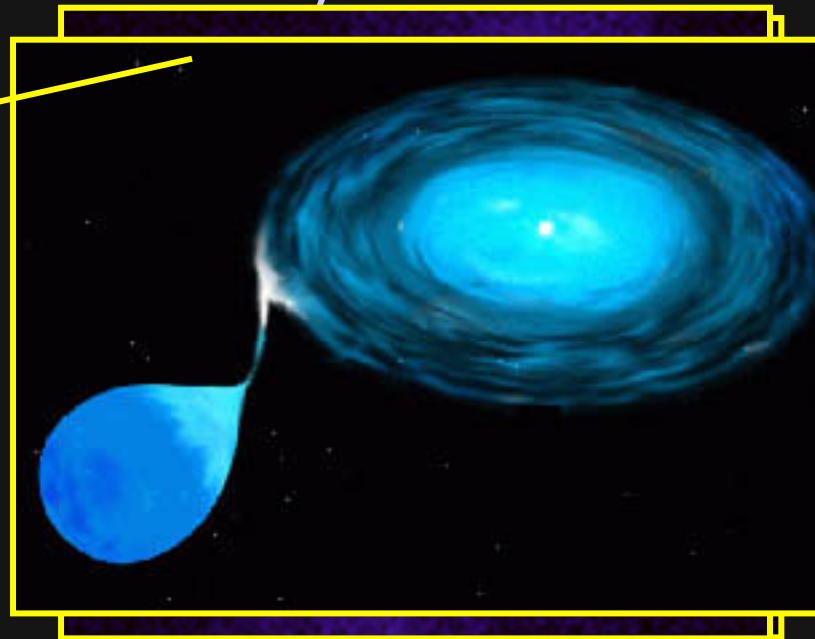


“Invisible” Light from Galaxies



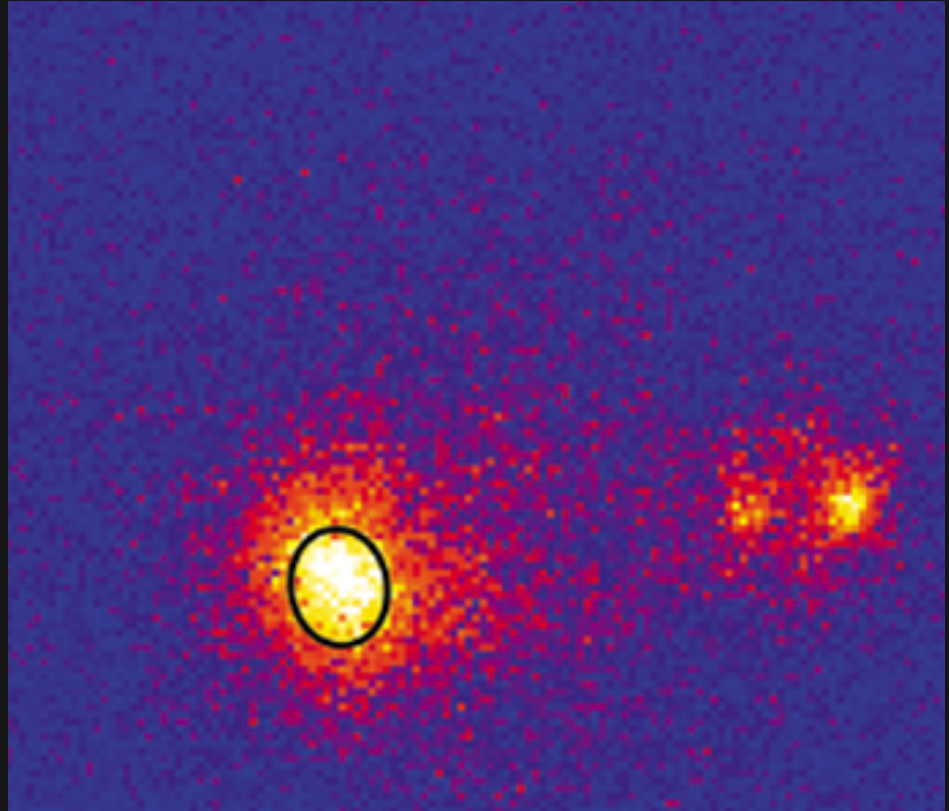
X-ray Objects in Galaxies:

- stars
- supernova remnants
- X-ray binaries



Hidden Mass in Galaxies

This X-ray image of an elliptical galaxy reveals hot, fast-moving gas even in the outer reaches of the galaxy. The visible mass of the galaxy is insufficient to hold onto it.

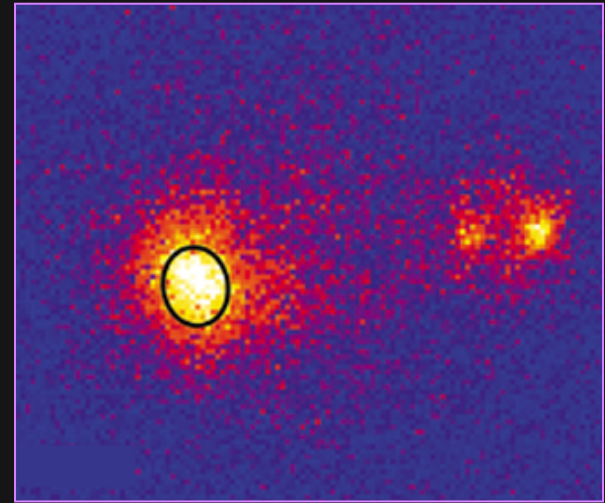


(The dark circle shows the size of the galaxy when photographed in visible light. The X-ray image shows mass far outside the visible image.)

Hidden Mass in Galaxies

So, we have a Problem:

- This gas gives off X-rays, which means it's *hot!*
- Hot gas moves at high velocities - we can measure and confirm this
- The velocity of the gas is greater than the escape velocity of the galaxy, if we calculate the galaxy's mass by adding up all the mass we can see at all wavelengths of light
- **So why hasn't the gas escaped? There *must* be more mass we *can't* see!**



Hidden Mass in Galaxies

Another way to look at the problem:

We can determine the mass of an object by measuring the motion of bodies in orbit around it.

Newton's Second Law:

$$F = ma$$

F = Force of Gravity
a = acceleration due
to circular motion

$$\frac{GMm}{r^2} = \frac{mv^2}{r}$$

Hidden Mass in Galaxies

From previous slide, we have Newton's Second Law of Motion assuming a gravitational force and acceleration due to circular motion:

$$\frac{GMm}{r^2} = \frac{mv^2}{r}$$

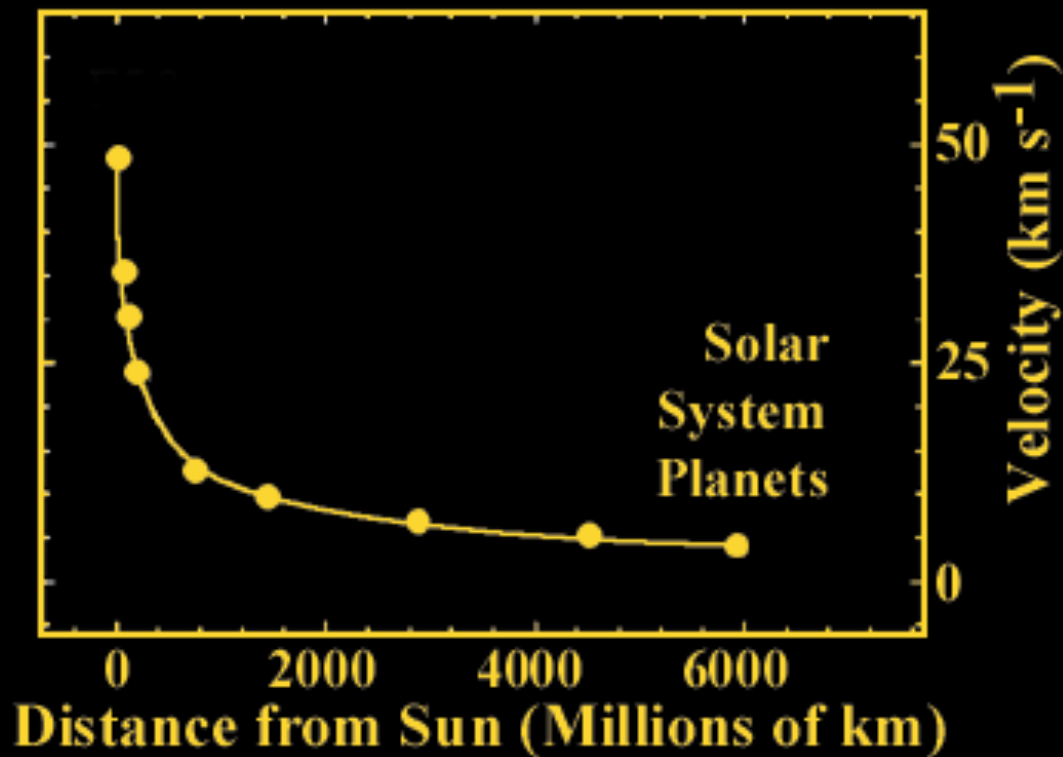
Simplifying gives:

$$v^2 = \frac{GM}{r}$$

So, if GM is constant, then velocity is proportional to the inverse square root of distance. For example ...

Hidden Mass in Galaxies

Rotation Curve - A Plot of Object Velocity vs. Distance



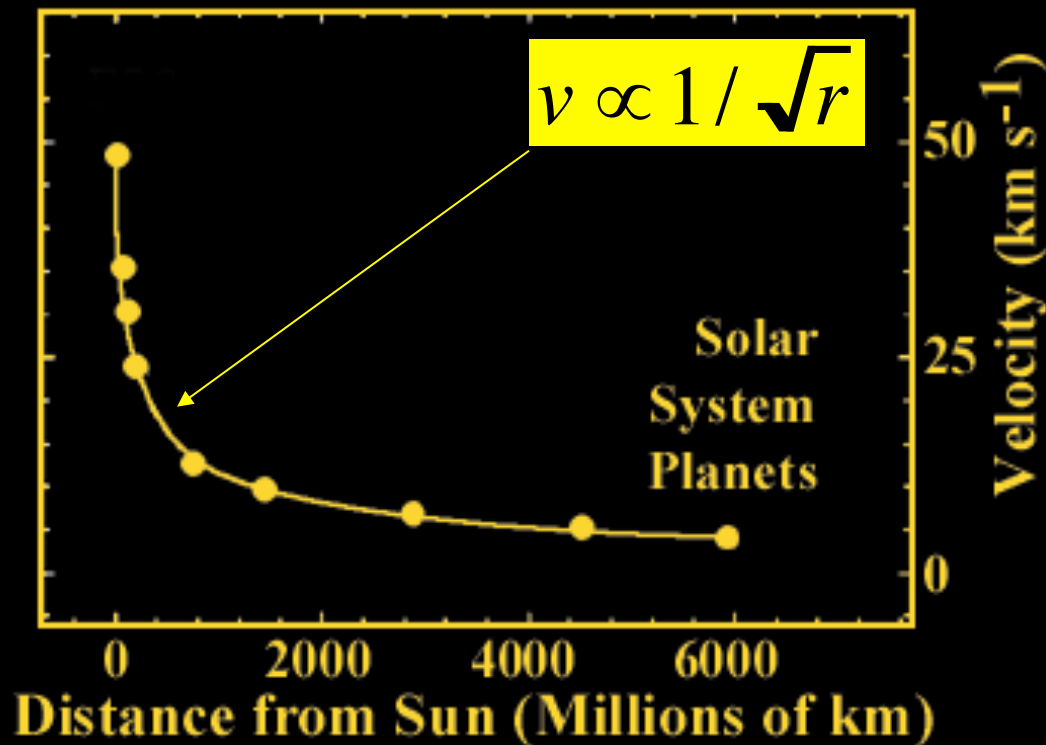
Activity #6a: Evidence for Hidden Mass

There are nine solar system planets presented on the graph. The planets, from the closest to the sun to the furthest, are Mercury, Venus, Earth, Mars, Jupiter, Saturn, Uranus, Neptune, and Pluto . Using the graph, the velocities of the solar system planets, from the lowest value to the highest value, are approximately 48, 35, 30, 24, 13, 10, 7, 5, and 4 km/sec . Using the graph, the distances of the planets from the Sun are, from least to greatest, 0, 110, 150, 250, 800, 1500, 2800, 4500, and 6000 million km . In general, the further a planet is from the sun the slower its velocity. The closer a planet is to the sun faster its velocity.

Hidden Mass in Galaxies

Rotation Curve - A Plot of Object Velocity vs. Distance

Our Solar System



Hidden Mass in Galaxies

The previous plot showed data for planets in our solar system, illustrating the equation:

$$v^2 = \frac{GM}{r}$$

If we solve for M:

$$M = v^2 r / G$$

We can use real data (the distances and velocities of the planets) and the fact that $G = 6.67 \times 10^{-11} \text{ m}^3/\text{kg}\cdot\text{s}^2$ to verify that the central mass, M , remains constant..

Activity 6b, part 1

Planet	Distance from Sun (km)	Velocity (km/s)	Mass (kg)
Earth	1.5×10^8	29.8	2.0×10^{30}
Jupiter	7.8×10^8	13.1	2.0×10^{30}
Neptune	4.5×10^9	5.4	2.0×10^{30}

Hidden Mass in Galaxies

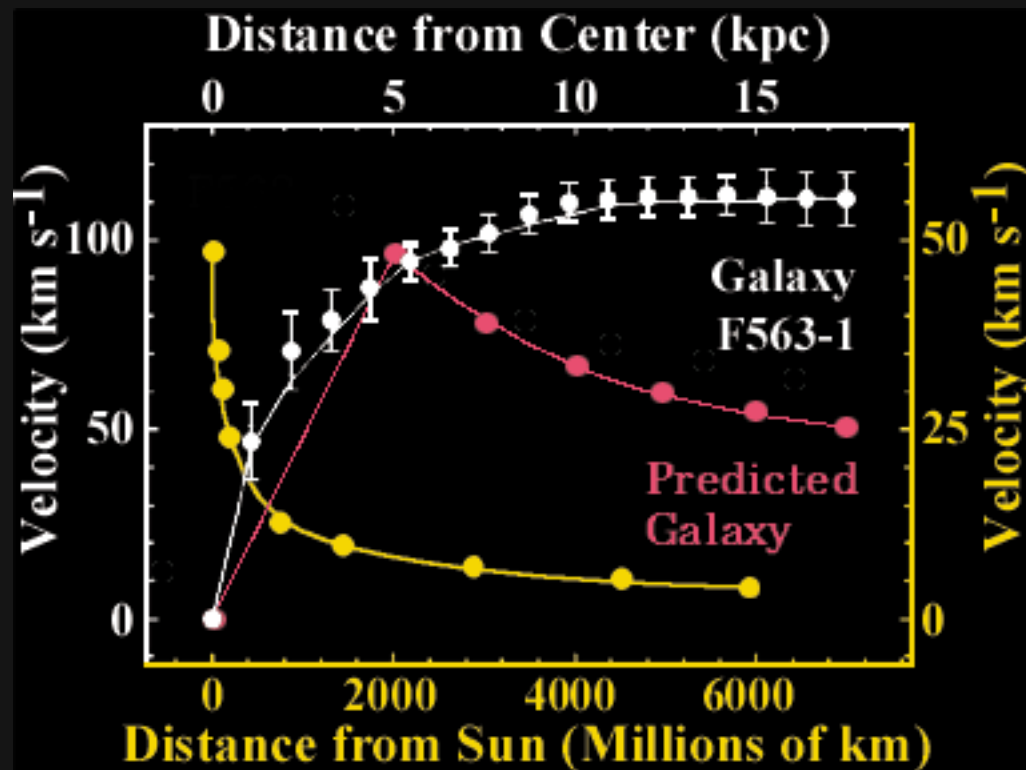
What does this have to do with Galaxies???

If we measure the distance and velocity of objects (eg., stars) orbiting in a galaxy, we'd expect them to obey the same laws. As distance from the center of the galaxy increases, we should get to a point where almost all of the galaxy's mass is inside the orbit of the furthest objects.

At this point, the central mass would be practically constant, and we would expect a rotation curve similar to that for our solar system.

Hidden Mass in Galaxies

What we expect for a galaxy if all the mass was concentrated in the central region.
Compare expected velocities with actual velocities



Hidden Mass in Galaxies

If we take the same equation we used for the solar system:

$$M = v^2 r / G$$

and use the actual distances and velocities observed for this galaxy, we can calculate the enclosed mass at various distances from the galaxy's center.

(Remember that $G = 6.67 \times 10^{-11} \text{ m}^3/\text{kg}\cdot\text{s}^2$)

Activity 6b, part 2

Distance (kpc)	Velocity (km/s)	Mass (kg)
5.0	95.0	2.1×10^{40}
10.0	110.0	5.6×10^{40}
15.0	110.0	8.4×10^{40}

Hidden Mass in Galaxies

What Happened?!?!?

The central mass never becomes constant, as it did for the Solar System. The fact that the mass is still increasing means we haven't yet reached a distance where all the mass is contained inside that orbit. But we've plotted *all the matter we see!*

There *must* be Missing Mass!!

Hidden Mass in Galaxies

Hot Gas and Rotation Curves show:

- Gas and objects move at velocities greater than can be accounted for from the gravitational effect of the visible mass of the galaxy.
- From these observations we deduce that the visible mass accounts for only 10 % of the total mass of the galaxy.
- **Recent results from WMAP show ordinary matter makes up only 4% of the universe!**

Candidates for the Hidden Mass

- Hydrogen Gas
 - Very abundant, but not enough detected
- MACHOs (Massive Compact Halo Objects)
 - E.g. Black Holes, Neutron Stars, Brown Dwarfs
 - Not enough of them
- WIMPs (Weakly Interacting Massive Particles)
 - E.g. Exotic subatomic particles
 - The best candidate theoretically, but not yet observed.

